

Large Synoptic Survey Telescope (LSST) Data Management

# **LSST Alerts: Key Numbers**

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Abstract

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# **LSST Alerts: Key Numbers**

**Alerts Key Numbers** 

# **1** Introduction

This document is currently in development and contains a lot of detailed information for the purposes of internal discussion; to be reduced/clarified later. Text in red needs some discussion by DM-SST.

The LSST Data Management System's (DMS) Alert Production (AP) pipeline will process new data as it is obtained by the telescope. Difference Imaging Analysis (DIA) will be performed, and all sources with a signal-to-noise ratio transSNR > 5 in positive or negative flux will be considered "detected", instantiate a record in the source catalogs, and generate an alert (LSR-REQ-0101 LSE-29; DMS-REQ-0269, -0274 LSE-61). Each alert is a packet containing LSST data about the source such as coordinates, photometry, and image cutouts. For a full description of detected sources and alert packet contents, see LSE-163. The LSST alert stream will be delivered to several community-developed brokers, and also accessible to users via the LSST Science Platform's alert filtering service (AFS); plans and policies for alert distribution are provided in LDM-612.

The purpose of this document is to quantitatively inform broker developers, and the broader scientific community planning to use alerts, on the key numbers regarding alert generation, distribution, and access via the LSST alert filtering service. The goals of this document are threefold: (1) to provide all of the key numbers regarding alert generation in one place; (2) to include any and all basis information, assumptions, and derivations that contributed to the key number; and (3) to be clear about whether each key number represents an estimate, a requirement, or a boundary.

In this work we use 8 bits per byte (B), and 1024 B per KB, 1024 KB per MB, and so forth.

A list of resources being used in the preparation of this document.

- LSST: From Science Drivers to Reference Design and Anticipated Data Products, Ivezic et al. (2008)
- LSST Science Requirements Document (SRD), LPM-17.



- LSST System Requirements (LSR), LSE-29.
- Observatory System Specifications (OSS) document, LSE-30.
- Data Management System Requirements (DMSR) document, LSE-61.
- Science Requirements and System Specifications Spreadsheet (SR&SSS), LSE-81.
- Data Products Definitions Document (DPDD), LSE-163
- Plans and Policies for LSST Alert Distribution, LDM-612
- Data Management Science Pipelines Design, LDM-151

# 2 Alert Stream

The concept and existence of the LSST alert stream is baked into the highest-level document: the SRD specifies that information about the detections of transient, variable, and moving objects be released as a data stream within 1 minute.

### 2.1 Alert Release Timescale

It is a formal requirement that the Data Management System release 98% of alerts for each visit within 60 seconds of the end of image readout.

**Formal Requirements** – Regarding the DMS's ability to generate alerts, the SRD states that *"data on likely optical transients ... will be released with a latency of at most* 0TT1 *minutes"*, where 0TT1 is 1/2/0.5 minutes (design/minimum specifications and stretch goal; LPM-17). The SRD's design specification value of 0TT1 flows down to a formal requirement in the LSR that specifies *"LSST shall meet the following specification for reporting of data on optical transients detected in single-visit data [in]* 0TT1", and defines 0TT1 as *"the latency of reporting optical transients following the completion of readout of the last image of a visit"* (LSR-REQ-0101, LSE-29). The LSR also makes a formal requirement that the DMS's *"algorithm shall be applied and the alert transmitted within the specified latency for at least a fraction* 0TR1 *of instances ... [and] remaining transients ... be identified and recorded at the next processing opportunity"*, where 0TR1 is 98% (LSR-REQ-0025, LSE-29). The LSR requirement also flows down to more specific formal requirement in the OSS and DMSR, which both state that *"Alerts shall be made available within time* 0TT1 from

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the conclusion of readout of the raw exposures used to generate each alert to the distribution of the alert to community distribution mechanisms" (OSS-REQ-0127, LSE-30; DMS-REQ-0004, LSE-61).

**Alerts Key Numbers** 

There is a mild amount of **tension in the formal requirements**: the SRD states that 0TT1 is a requirement on the *releasing* of an alert; the LSR states that 0TT1 is a requirement on the *reporting* of an alert (LSR-REQ-0101), and then later that it is the time for *transmitting* the alert to the broker (LSR-REQ-0025). It is only the OSS and DMSR which interpret 0TT1 as ending at the time of *distribution to the broker*. In this respect, **the OSS and DMSR are incorrect**: the time it will take for an alert to be transmitted to the broker is independent of the LSST Prompt Processing system (see § 2.5), and 0TT1 should end at the time the alert is *available to the stream*. A further tension is that whereas 0TT1 is flowed down from the LSR to the OSS and DMSR, 0TR1 is not (see further discussion in § 2.7).

### Potential Change to the OSS and DMSR:

OSS-REQ-0127 [LSE-30] and DMS-REQ-0004 [LSE-61] could have this statement updated from: "Alerts shall be made available within time OTT1 from the conclusion of readout of the raw exposures used to generate each alert to the distribution of the alert to community distribution mechanisms."

to instead say:

"Alerts shall be made available within time OTT1 from the conclusion of readout of the raw exposures used to generate each alert. The acceptable fraction of alerts released within a latency of OTT1 is OTR1."

In the DMSR, 0TR1 could be added to the table in Section 1.3.1 (*"Nightly Data Accessible Within 24 hrs"*), which also contains 0TT1. In the OSS, 0TT1 is used but not defined, so presumably 0TR1 could be used but not defined, too. Another 0TR1-related change to the OSS is proposed in § 2.7.

**Additional Notes –** We often colloquially say *OTT1* starts *at the time of shutter close*, but it actually starts 2 seconds later *at the end of readout* (or longer if slower reads are adopted).

**Requirement Breach Protocol** – For the OTR1 = 2% of alerts allowed to be transmitted with > OTT1 seconds, how will they be identified as delayed? I.e., with a flag in the alert packet?

Add some Science Drivers? What are the main science driveres behind OTT1 = 60 seconds? If OTT1 is descoped to, e.g., 2 minutes, what is the science impact? The SRD states *"Time scales ranging from*  $\sim 1 min$  (to constrain the properties of fast faint transients such as those recently



*discovered by the Deep Lens Survey)..."* are required for science – check what these fast faint transients are.

### 2.2 Number of Alerts per Visit

It is a formal requirement that the Data Management System can sustain the generation of an average of at least 10000 alerts per visit.

**Formal Requirements** – Regarding the DMS's ability to generate alerts, the SRD states that *"The system should be capable of reporting such data for at least* transN *candidate transients per field of view and visit,"* where transN is  $10^4/10^3/10^5$  (design/minimum specifications and stretch goal; LPM-17). The design specification value of transN flows down to a formal requirement in the LSR, which describes transN as *"the minimum number of optical transients for which data can be reported per visit"*, with a note that *"it is unclear whether the SRD specification of* transN *refers to the number of alerts that can be generated for a single visit (i.e. an instantaneous limit), or the number per visit averaged over time"* (LSR-REQ-0101, LSE-29). The OSS does clarify: *"The LSST Data Management system shall be sized to accommodate an average value of at least* nAlertVisitAvg alerts generated per standard visit while meeting all its other requirements," where nAlertVisitAvg is  $10^4$  (OSS-REQ-0193 in LSE-30). No corresponding requirement is flowed down to the DMSR, but the SR&SSS also uses the same minimum average value, and furthermore uses 40000 as the *peak* number of alerts per visit, with a note that it was chosen as a number in between the SRD's minimum of  $10^4$  and the stretch goal of  $10^5$  [LSE-81].

There is a mild amount of **tension in the formal requirements**, in that neither transN nor nAlertVisitAvg are flowed down to the DMSR.

### Potential Change to the DMSR:

If nAlertVisitAvg (which is essentially equivalent to transN, but more specific) should exist in the DMSR, then perhaps a new sub-section should be added to Section 2.2.3 *"Transient Alert Distribution"* such as:

Section 2.2.3.1 "Performance Requirements for Transient Alert Distribution"

ID: DMS-REQ-XXXX

Specification: The system shall be able to identify and release an average of at least nAlertVisitAvg alerts per visit.

Table: "Average minimum number of alerts per visit required to be supported.", 10000, alerts,



nAlertVisitAvg.

Derived from Requirements: LSR-REQ-0101, OSS-REQ-0193

**Science Drivers** – The value of transN = 10000 alerts per visit is a formal requirement on the DMS, not a scientific estimate of the intrinsic rate of transients and variables in the universe. The number of alerts is expected to be lower/higher in extra/galactic fields. As derived from the Science Book (LSST Science Collaboration, 2009), the estimates for the most common transients and variables are as follows:

- Variable Stars: LSST is predicted to observe a total of  $\sim 135$  million variable stars. Making the simple assumption that 20/80% are in extra/galactic fields, and that of the  $\sim 18000$  $deg^2$  surveyed by LSST 80/20% of the fields are extra/galactic, and that 10% of all variable stars are detectably variable at any given time, then a typical extra/galactic field would yield  $\sim 1800/28800$  alerts per visit. Averaged over all fields, and weighted by 80/20% of the fields being extra/galactic, this is 7200 alerts per visit.
- Supernovae: LSST is predicted to observe a total of 10 million supernovae in 10 years, or 1 million per year. Since SNe are typically only visible for a few months, there might be  $\sim 0.3$  million detectable at any given time. Over  $15000 \ deg^2$  of extragalactic survey area, that's ~  $20 \text{ SNe deg}^{-2}$  or ~ 200 SNe per visit.
- Active Galactic Nuclei: LSST is predicted to observe millions of AGN. If  $\sim 10\%$  of them are detectably variable at any given time, then the estimate is  $\sim 0.1$  million alerts over  $15000 \text{ deg}^2$  would generate ~ 7 alerts  $\text{deg}^{-2}$  or ~ 70 alerts per visit for AGN.
- Moving Objects: The predicted number of Solar System objects that LSST is predicted to observe is dominated by the 5.5 million main-belt asteroids. Assuming that they are spread evenly over the  $\sim 18000 \text{ deg}^2$  survey area (they're not, as they're found primarily along the ecliptic) leads to  $\sim 3000$  alerts per visit due to moving objects.

Therefore, astrophysical estimates for the occurrence rates of alerts caused by the most common types of transients and variables yield  $\sim 5100/32000$  alerts per visit in extra/galactic fields, with an average of  $\sim 10500$  alerts per visit.

Requirement Breach Protocol – Is there a new upper limit on the peak volume? In cases where there are > 40000 alerts generated by a visit, does DM expect they issued with a delay?

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### 2.3 Alert Packet Size

There are no formal requirements regarding the alert packet size. The size of an individual alert packet is estimated to be  $\leq 82$  KB (without schema or compression). Eric is working on improved alert packet simulations for a more accurate sizing estimate.

**Sizing Estimate** – Alert packet contents will include all of the LSST science data for the triggering detection, including a ~ 12 month historical record of detections, plus image stamp cutouts. Alert contents are described in more detail in Section 3.5 of LSE-163. Simulated alert packets based on the Apache Avro format are at most ~ 82/126 KB, without/with the schema, respectively. This volume represents an alert packet for a variable star with a full 12 month history of detections. For example, a new unassociated source with only a single detection would be ~? KB, and a ~ 1 month long transient followed by ~ 11 months of forced photometry would be ~? KB. The application of gzip compression can further reduce the size of a ~ 82/126 KB alert to ~ 68/65 KB (JIRA ticket DM-16280). Cutout stamps included in the alert will be at least  $30 \times 30$  pixels and contain flux (32 bit/pix), variance (32 bit/pix), and mask (16 bit/pix) extensions for both the template and difference image, plus a header of metadata [LSE-163]. The stamps alone will contribute  $\geq 18$  KB to the total size of the uncompressed alert packet (i.e., ~ 20%).

Alternative "Lite" Content Options – Brokers which plan to do their own source association, compile source catalogs based on alerts, or not use the image stamps might prefer a stream of packets with appropriately reduced information. The LSST DM team currently expects that some options will be possible, and brokers may propose an option that works for them during the selection process [LDM-612]. As previously mentioned, removing the image stamps would reduce packet size by  $\geq 18$  KB. Removing the historical records of past detections would reduce all alert packets to be equivalent in size to a new unassociated source. A few of these options might also be available to users of the LSST alert filtering service (§ 3).

### 2.4 Alert Stream Data Rate

There are no formal requirements regarding the alert stream data rate. The *time-averaged* data rate of the alert stream is estimated to be  $\sim 27 \text{ MB/sec}$  (for alert packets without schema or compression), potentially with bursts of up to 640 MB/sec.



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**Sizing Estimate –** The size of a single LSST alert will be  $\sim 82$  KB (including image stamps but not schema nor compression). Using an average of  $\sim 10000$  alerts released per  $\sim 30$  second image, this leads to a *time-averaged* alert stream data rate of  $\sim 27 \text{ MB s}^{-1}$ . As discussed in § 2.2, the number of alerts per field will vary in extra/galactic fields from  $\sim 2000$  to  $\leq 40000$ , which would produce *time-averaged* alert streams of  $\sim 5.4$  to  $\leq 108$  MB s<sup>-1</sup>. However, in order to release alerts within 0TT1 = 60 seconds of image readout (§ 2.1), the stream will not be continuous in time, but periodic, with potential bursts. For example, if all 10000 alerts are issued within the last 5 seconds of 0TT1 this would produce a data rate of 160 MB s<sup>-1</sup>; in galactic fields with  $\leq 40000$  alerts this could be as high as  $640 \text{ MB s}^{-1}$ .

#### 2.5 Number of Selected Brokers

There are no formal requirements on the number of brokers. The DM team estimates that resources will allow for the delivery of the alert stream to 4-7 brokers.

Formal Requirements – Neither the SRD, LSR, OSS, nor the DMSR place any formal requirements on the number of brokers that the alert stream should be delivered to, and the DPDD and the SR&SSS do not mention an estimate for the number of brokers.

Sizing Estimate - As described in Section 2.2.3 of LDM-612, "An allocation of 10 Gbps is baselined for alert stream transfer from the LDF, with an estimated packet size of 82 KB and up to 10,000 alerts per visit. For illustration, based on these numbers up to 7 brokers could receive the full stream if 5 seconds is budgeted for outbound data transfer." Is there another document to cite for the allocation of 10 Gbps?

### 2.6 Alert Database Volume

There are no formal requirements on the alerts database volume. The estimated maximum upper limit is  $\leq 3 \text{ PB}$  (without schema or compression).

**Sizing Estimate –** An upper estimate is derived by starting with a maximum of  $\sim 1000$  visits per night, and  $\sim 10000$  alerts per visit, which amounts to  $\sim 10$  million alerts per night. Using 2000/36000 alerts per visit for extra/galactic fields and assuming 80/20% of the 1000 visits per night are for extra/galactic fields, as described in § 2.2, also generates  $\sim 10$  million alerts per night. Assuming the upper estimate of  $\sim 82$  KB per alert (§ 2.3), that leads to a total of



~ 782 GB per night. An extreme upper limit is 365 nights per year for 10 years, which would amount to ~ 2.7 PB at the very most. Therefore we quote an extreme upper limit on the alerts database as  $\leq$  3 PB. Compression could drastically lower this, as could reformatting: every alert contains a ~ 12 month historical record and links to the most recent DIAObject and DR Object catalogs. The set of alerts for the same transient/variable would contain significant redundant information which could be reformatted (i.e., removed and compiled).

### 2.7 Delayed/Failed Alert Distribution

It is a formal requirement that 98% of all alerts for a given visit are issued within 60 seconds.

It is a formal requirement that < 1% of all science visits have any fraction of their alerts experience a distribution delay > 60 seconds.

It is a formal requirement that < 0.1% of all science visits experience a total failure in alert generation and distribution.

**Formal Requirements** – The SRD does not say anything on the topic of alert distribution delays or failures. As mentioned in § 2.1, the LSR defines OTR1 as the *"fraction of detectable alerts for which an alert is actually transmitted within latency* OTT1", where OTR1 = 98% (LSR-REQ-0025; LSE-29). The OSS does not state any requirements on the fraction of failed alerts per visit, but does specify that *"no more than* sciVisitAlertFailure *% of science visits ... shall fail to be subjected to alert generation and distribution"*, where sciVisitAlertFailure = 0.1%, and that *"no more than* sciVisitAlertDelay *% of science visits ... shall have their alert generation and distribution completed later than* [OTT1]", where sciVisitAlertDelay = 1% (OSS-REQ-0112; LSE-30). The OSS furthermore makes the distinction that if any number of the alerts for a given visit are distributed later than OTT1, it counts towards *sciVisitAlertDelay*. The DMSR makes no statements about the fraction of alerts per visit with delayed/failed distribution, or the fraction of visits with failed/delayed alert distribution.

There is some **tension in the formal requirements**, as the OSS does not use OTR1 (= 98%, the acceptable fraction of alerts per visit released within OTT1) to define a failed/delayed visit. Instead, if a visit has any number of alerts delayed beyond OTT1 it is considered as counting towards sciVisitAlertDelay. Basically, LSR-REQ-0025 and OTR1 do not appear to have flowed down to the OSS or DMSR, and OTR1 does not seem to actually be used for anything (unless it's flowed down directly to science verification documents). Furthermore, there is no equiv-



alent to sciVisitAlertFailure and/or sciVisitAlertDelay in the DMSR. Finally, there is no specification for an acceptable fraction of alerts which completely fail to be distributed (i.e., a requirement like 0TR1, but for the fraction of alerts with a failed instead of a delayed release).

### Potential Change to the OSS:

The definition of sciVisitAlertDelay in the OSS could be updated with the blue text to include OTR1, such as:

Section 3.1.2.1 "Science Visit Alert Generation Reliability"

ID: OSS-REQ-0112

Specification: No more than sciVisitAlertDelay % of science visits read out in the camera [and specified to be analyzed by Data Management] shall have their alert generation and distribution completed later than the SRD specification for alert latency 0TT1; this applies to 0TR1 % of the alerts from that visit as specified in the LSR.

Note: OTT1 is used but not defined in the OSS, and OTR1 is not yet defined in the OSS either.

### Potential Change to the DMSR:

If sciVisitAlertFailure and/or sciVisitAlertDelay should be included as formal requirements in the DMSR, then perhaps a new sub-section could be added to Section 2.2.3 *"Transient Alert Distribution"* such as:

Section 2.2.3.1 "Alert Delay and Failure Tolerances"

ID: DMS-REQ-????

Specification: Of the science visits read out of the camera and specified to be analyzed by Data Management, no more than sciVisitAlertDelay = 1% and sciVisitAlertFailure = 0.1% shall have their alert generation and release delayed or failed, respectively (integrated over all stages of data handling).

Derived from Requirements: OSS-REQ-0112

### Potential Change to the LSR:

Should there be a specification which is like OTR1 (which specifies that = 98% of alerts per visit must not be delayed beyond OTT1) but specifies the acceptable fraction of alerts per visit which fail to be released? Should this just be 0%? Or 0.1% like sciVisitAlertFailure?

**Requirement Breach Protocol** – Are alerts from a delayed visit flagged in some way? With regards to OTR1 = 98%, LSR-REQ-0025 states that "*the remaining transients so detectable must still be identified and recorded at the next processing opportunity*", but this is not this flowed down to DMSR, and it is unclear what "*the next processing opportunity*" means.





### 2.8 Alert Stream Completeness and Purity

It is a formal requirement that DM derive and supply threshold values for a *spuriousness* parameter (also known as *real/bogus*), which can be used to filter all alerts into a subsample with a specified completeness and purity, and reduce the fraction of false positives per visit (i.e., sources detected that are not astrophysical in origin).

**Formal Requirements** – The SRD makes no statements about alert stream purity or completeness, but does quote that the *"minimum signal-to-noise ratio in difference image for reporting detection of a transient object"* has a design specification of transSNR = 5 [LPM-17]. The LSR contains essentially the same definition for transSNR, *"the signal-to-noise ratio in single-visit difference images above which all optical transients are to be reported"* (LSR-REQ-0101; LSE-29). There is no minimum specification or stretch goal associated with transSNR. However, as described in the DPDD, detections with transSNR > 5 but *"with high probability of being instrumental non-astrophysical artifacts may be excluded"* (Section 3.1, LSE-163), and a limited number of transSNR < 5 sources might be allowed to generate alerts (e.g., sources near a gravitational keyhole which could make an asteroid potentially hazardous; Section 3.2.1, LSE-163).

Regarding the alert stream completeness and purity, the OSS requires that "there shall exist a spuriousness threshold T for which the completeness and purity of selected difference sources are higher than transCompletenessMin and transPurityMin, respectively, at the SNR detection threshold transSampleSNR. This requirement is to be interpreted as an average over the entire survey" (OSS-REQ-0353; LSE-30). In other words, the DMS must be able to provide the value for a spuriousness threshold T, below which all alerts detected with a signal-to-noise ratio transSampleSNR = 6, over the entire LSST survey, have transCompletenessMin = 90% and transPurityMin = 95%. (For moving objects there is a separate specification for a threshold pertaining to transSampleSNR = 5, mopsCompletenessMin = 99%, and mopsPurityMin = 50%; OSS-REQ-0354, LSE-30). This spuriousness threshold T will allow users to filter their stream to a fiducial completeness and purity; brokers may request a pre-filtered stream that includes a restriction on spuriousness.

# 3 The LSST Alert Filtering Service

It is a formal requirement that the LSST provide a simple alerts filtering service for users (individuals with LSST data rights and access to the Science Platform), which is hereafter referred





to as the LSST alert filtering service (AFS).

**Formal Requirements** – The SRD specifies that "users will have an option of a query-like prefiltering of [the alert] data stream in order to select likely candidates for specific transient type" and that "several pre-defined filters optimized for traditionally popular transients, such as supernovae and microlensed sources, will also be available" [LPM-17]. Neither the LSR nor the OSS have a formal requirement on this capability, as it is a product of the DMS. The DMSR has a formal requirement that "a basic, limited capacity, alert filtering service shall be provided that can be given user defined filters to reduce the alert stream to manageable levels", and that this service include "a predefined set of simple filters" (DMS-REQ-0342, -0348; LSE-61).

### 3.1 Number of Simultaneous AFS Users

It is a formal requirement that the AFS support a minimum of 100 simultaneous users.

**Formal Requirements** – The DMSR specifies that the LSST *"alert filtering service shall support* numBrokerUsers *simultaneous users*", where numBrokerUsers = 100 (DMS-REQ-0343; LSE-61).

### 3.2 Number of Alerts per Visit Returned per User-Defined Filter

It is a formal requirement that the AFS return 20 alerts per visit per user.

**Formal Requirements** – The DMSR specifies that within the LSST alert filtering service *"each user [shall be] allocated a bandwidth capable of receiving the equivalent of* numBrokerAlerts *alerts per visit"*, where numBrokerAlerts = 20 (DMS-REQ-0343; LSE-61).

Note that in a footnote of LDM-612, it says that the *"requirement on the number of simulta-neously connected users and number of passed alerts is largely driven by outbound bandwidth limitations from the DAC at NCSA. We are investigating approaches that would support larger numbers of active filters"* (page 12; LDM-612).

### 3.3 Alerts Database Query Latency

It is a formal requirement that all alerts be stored in a database and available for query (by users with data rights and access to the science platform).



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**Formal Requirements** – The OSS states that *"All published transient alerts ... shall be available for query"* (OSS-REQ-0185; LSE-30). Like all other Prompt data products, the Alerts Database will be updated within L1PublicT = 24 hours (LSR-REQ-0104, LSE-29).



## References

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